

Record Ultra-low Phase Noise 12 GHz Signal Generation with a Fiber Optical Frequency Comb and Measurement

Xiaopeng Xie, Romain Bouchand, Daniele Nicolodi, Michele Giunta, Wolfgang Hänsel, Matthias Lezius, Ronald Holzwarth, Abhay Joshi, Shubhashish Datta, Christophe Alexandre, et al.

▶ To cite this version:

Xiaopeng Xie, Romain Bouchand, Daniele Nicolodi, Michele Giunta, Wolfgang Hänsel, et al.. Record Ultra-low Phase Noise 12 GHz Signal Generation with a Fiber Optical Frequency Comb and Measurement. CLEO: Science and Innovations, 2016, San Jose, United States. pp.SM4H.1, 10.1364/CLEO_SI.2016.SM4H.1. hal-02468267

HAL Id: hal-02468267 https://cnam.hal.science/hal-02468267v1

Submitted on 27 Mar 2020

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

Record Ultra-low Phase Noise 12 GHz Signal Generation with a Fiber Optical Frequency Comb and Measurement

Xiaopeng Xie¹, Romain Bouchand¹, Daniele Nicolodi¹, Michele Giunta², Wolfgang Hänsel², Matthias Lezius², Ronald Holzwarth², Abhay Joshi³, Shubhashish Datta³, Christophe Alexandre⁴, Pierre-Alain Tremblin⁵, Giorgio Santarelli^{1,5}, and Yann Le Coq^{1*}

Giorgio Santarelli^{1,5}, and Yann Le Coq^{1*}

¹LNE-SYRTE, Observatoire de Paris, CNRS, UPMC, 61 avenue de l'Observatoire, 75014 Paris, France.

² Menlo Systems GmbH, Am Klopferspitz 19a, D-82152 Martinsried, Germany

³Discovery Semiconductors Inc., 119 Silvia street, Ewing, New Jersery 08628, USA

⁴LNE-CNAM, CEDRIC Laboratory, 1 rue Gaston Boissier, 75015 Paris, France

⁵Laboratoire Photonique, Numérique et Nanosciences, IOGS - CNRS - Université de Bordeaux1, Rue F. Mitterand, 33400 Talence, France

* E-mail address: yann.lecoq@obspm.fr

Abstract: We demonstrate a 12 GHz signal with a record absolute phase noise of -167 dBc/Hz at 10 kHz and -170 dBc/Hz at 100 kHz. We have developed a specific measurement setup for charactering this ultra-low phase noise.

OCIS codes: (140.4050) Mode-locked laser; (230.5160) Photodetectors; (350.4010) Microwave; (120.5050) Phase measurement.

1. Introduction

Microwave signals with ultra-low phase noise are of great interested in the fields of radar, telecommunication, deep-space navigation, very long baseline interferometry and precision metrology [1]. Especially, the microwave signals extracted from the optical frequency comb possess extremely low phase noise not only at the close-to-carrier Fourier frequencies but also at the far-from-carrier Fourier frequencies [2-6].

In this contribution, we will report on our use of a fiber optical frequency comb locked to a CW ultra-stable laser to generate a 12 GHz carrier microwave signal with record low absolute phase noise levels of -167 dBc/Hz at 10 kHz and -170 dBc/Hz at 100 kHz Fourier frequencies. This absolute phase noise is 10 dB lower than all the previously reported data using a similar approach [2-6]. Furthermore, this phase noise level is also comparable or even lower than that of signals generated by the best sapphire whispering-gallery microwave oscillator. Measuring such ultra-low phase noise is in itself an interesting challenge. We have setup for this purpose a heterodyne cross-correlation technique with ultra-low measurement noise that uses two accessory statistically independent microwave sources as references.

2. Principle and experimental setup

Fig.1 shows the ultra-low phase noise 12 GHz signal generation setup. A fast seft-starting Er-doped fiber-based optical frequency comb with 250 MHz repetition rate is phase locked to an ultra-stable 1542 nm CW laser reference. A specially designed highly linear photodiode is used to convert the locked optical pulses train to electrical signals, generating harmonics of the comb repetition rate. The frequency comb thus transfers the spectral purity of the state-of-art ultra-stable CW laser to the microwave domain.

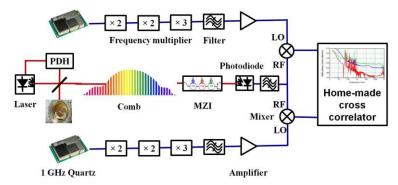


Fig. 1: Schematic of the ultra-low phase noise 12 GHs signal generation system and phase noise measurement set-up.

Several key technologies are applied to get the record ultra-low phase noise result. An external repetition rate multiplier redistributes the photocurrent to the harmonics at 12GHz, greatly increasing the signal to noise limit [7]. Furthermore, as the shot noise limit depends on the optical pulse duration [8], a piece of dispersion compensated

fiber is introduced to compress optical pulse on the photodiode to less than 1ps. As the comb's amplitude noise produces excess phase noise through amplitude phase conversion in the photodetector, we need to decrease this effect by 1) decreasing the combs' amplitude noise with passive optimization combined with an active servo loop 2) operate the photodetector at a "magic point" where non-linear saturation effect exactly cancel the amplitude-to-phase conversion noise for the 12 GHz signal [9,10]. Rejection of amplitude noise by more than 30dB and up to 50 are typically obtained and used.

To measure the resulting ultra-low microwave phase noise, we use a heterodyne cross-correlation technique, where two extra statistically independent microwave sources are introduced to implement a three cornered hat comparison (Fig.1). These two extra 12 GHz reference sources are obtained by multiplication of commercial 1 GHz oscillators. The two beat-note signals (near 10 MHz) obtained by beating these sources against the comb-generated signal are sent to a FPGA-based home-made system. They are sampled by fast analog-to-digital converters, digitally down-converted and computer processed to generate two time dependant phase comparison data sets. Cross-correlation of the these two phase noise data sets reveals the power spectral density of the phase noise of the 12 GHz signal that we want to characterize. The uncorrelated noise from the two accessory reference sources is rejected and contributes only to the uncertainty of the estimates of the phase noise power spectra density of the signal under test. This uncertainly averages down with the square root of the inverse of the measurement time.

3. Results and discussion

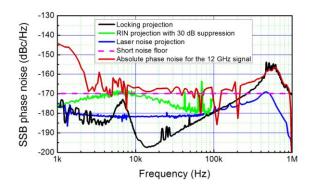


Fig. 2: Absolute SSB phase noise for the 12 GHz signal extracted from the fiber frequency comb and limitations analysis.

Fig.2 shows the absolute phase noise power spectral density measured by the home-made cross-correlator after several hours' average. The red curve represents the absolute phase noise of the 12 GHz signal generated with the fiber optical frequency comb. Other data sets, using frequency comb-based systems as auxiliary sources also show phase noise at 1Hz Fourier frequency below -100dBc/Hz. Blue, black and green traces show some identified contribution to this phase noise. The absolute phase noise at the 100 kHz Fourier frequency is limited by the shot noise floor, which we plan to improve in the near future by using higher power fs-laser combined with higher power handling photodiodes. Ultra-low phase noise -180dBc/Hz at 100 kHz seams within reach for the next generation system currently under development.

The authors would like to thanks Dawei Li, Jose Pinto and Michel Lours for the frequency chain implementation. This work is funded by the DARPA PULSE program (PureComb project), The FIRST-TF Labex and the Eurostar Eureka program (STAMIDOF project).

4. References

- [1] J. Millo, M. Abgrall, M. Lours, E. M. L. English, H. Jiang, J. Guéna, A. Clairon, S. Bize, Y. Le Coq, and G. Santarelli, Appl. Phys. Lett. 94, 141105 (2009).
- [2] A. Bartels, S. A. Diddams, C. W. Oates, G. Wilpers, J. C. Bergquist, W. H. Oskay, and L. Hollberg, Opt. Lett. 30, 667–669 (2005).
- [3] J. Millo, R. Boudot, M. Lours, P. Y. Bourgeois, A. N. Luiten, Y. Le Coq, Y. Kersalé, and G. Santarelli, Opt. Lett. 34, 3707–3709 (2009).
- [4] T. M. Fortier, M. S. Kirchner, F. Quinlan, J. Taylor, J. C. Bergquist, T. Rosenband, N. Lemke, A. Ludlow, Y. Jiang, C. W. Oates, and S. A. Diddams, Nat. Photonics 5, 425–429 (2011).
- [5] A. Didier, J. Millo, S. Grop, B. Dubois, E. Bigler, E. Rubiola, C. Lacroûte, and Y. Kersalé, Appl. Opt. 54, 3682-3686 (2015)
- [6] E. P. Campa, G. Buchs, S. Kundermann, L. Balet, and S. Lecomte, Opt. Express 23, 32441-32451 (2015)
- [7] A. Haboucha, W. Zhang, T. Li, M. Lours, A. N. Luiten, Y. Le Coq, and G. Santarelli, Opt. Lett. 36, 3654–3656 (2011).
- [8] F. Quinlan, T. M. Fortier, H. Jiang, A. Hati, C. Nelson, Y. Fu, J. C. Campbell, and S. A. Diddams, Nat. Photonics 7, 290–293 (2013).
- [9] W. Zhang, T. Li, M. Lours, S. Seidelin, G. Santarelli, and Y. Le Coq, Appl. Phys. B 106, 301–308 (2012).
- [10] W. Zhang, S. Seidelin, A. Joshi, S. Datta, G. Santarelli, and Y. Le Coq, Opt. Lett. 39, 1204-1207 (2014)